# Fundamentals of Imaging Lenses

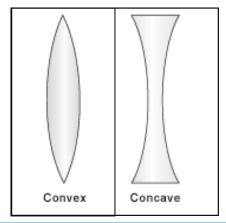
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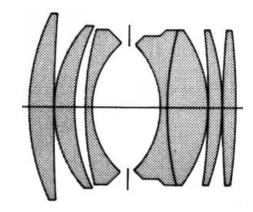
## This slide pack

In this part, we will introduce lenses

#### Lenses

- The components of an optical system consist of
  - aperture ring
  - Refractive elements
- Lenses:
  - Simple (left): single element
     Characteristics:
    - Refraction index
    - Shape of front+back
    - Often coated to improve optical properties
  - Compound (right): multiple lenses





#### Lenses

- Surface shapes:
  - Planar
  - Spherical
  - Aspherical: some surface which is not a sphere
- Call
  - d<sub>0</sub>: radius surface facing object plane
  - d₁: radius surface facing image plane
- Depending on positive or negative radius, one can have different single lens types



 $\begin{aligned} & \text{Bi-convex} \\ & d_0 > 0 \\ & d_1 < 0 \end{aligned}$ 



Bi-concave  $d_0 < 0$  $d_1 > 0$ 



Planar convex  $d_0 = \infty$   $d_1 < 0$ 



Planar concave  $d_0 = \infty$   $d_1 > 0$ 



Meniscus convex  $d_0 > 0$   $d_1 > 0$ 

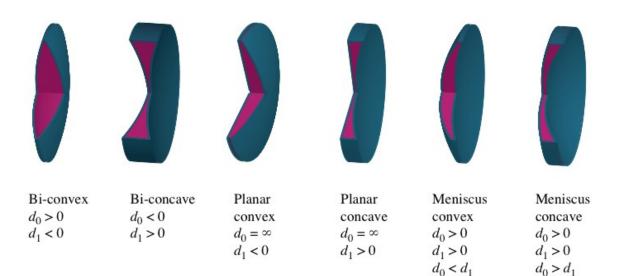
 $d_0 < d_1$ 



Meniscus concave  $d_0 > 0$  $d_1 > 0$  $d_0 > d_1$ 

#### Lenses

- Convex lenses direct light towards the optical axis: convergent or positive
- Concave lenses do the opposite and are called divergent or negative



5/21/19

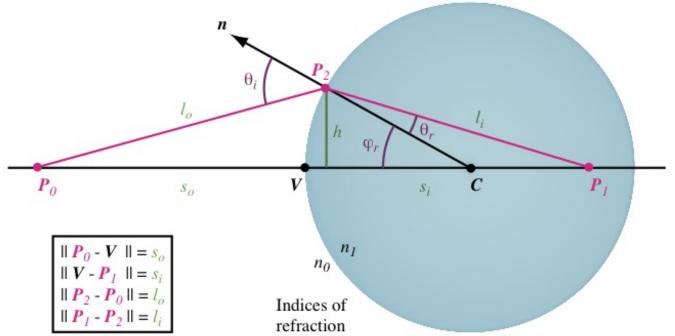
## **Spherical surface**

- Fermat's principle ⇒ optical path length of ray is n<sub>0</sub>l<sub>0</sub>+n<sub>1</sub>l<sub>i</sub>
- Path length of I<sub>o</sub> and I<sub>i</sub> is

$$l_o = \sqrt{d^2 + (s_o + d)^2 - 2d(s_o + d)\cos(\phi)},$$
  
$$l_i = \sqrt{d^2 + (s_i - d)^2 + 2d(s_i - d)\cos(\phi)}.$$

Substituting in the 1st

$$n_0 \sqrt{d^2 + (s_o + d)^2 - 2d(s_o + d)\cos(\phi)} + n_1 \sqrt{d^2 + (s_i - d)^2 + 2d(s_i - d)\cos(\phi)}.$$



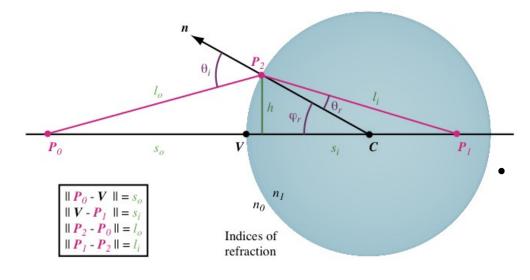
## Spherical surface

 Applying Fermat's principle by using Φ as the position variable:

$$\frac{n_0 d \left(s_o + d\right) \sin \left(\phi\right)}{2l_o} - \frac{n_1 d \left(s_i - d\right) \sin \left(\phi\right)}{2l_i} = 0.$$

thus: 
$$\frac{n_0}{l_o} + \frac{n_1}{l_i} = \frac{1}{d} \left( \frac{n_1 s_i}{l_i} - \frac{n_0 s_o}{l_o} \right)$$

 Rays from P<sub>0</sub> to P<sub>1</sub> with one refraction obey this law



- Remember, if the angle is too flat, refraction turns into reflection
- Under the hypotheses of Gaussian optics cos(Φ)≈1.
- If we consider only paraxial rays

$$l_o \approx s_o$$

$$l_i \approx s_i$$
.

thus the eq. on the left becomes

$$\frac{n_0}{s_o} + \frac{n_1}{s_i} = \frac{n_1 - n_0}{d}$$

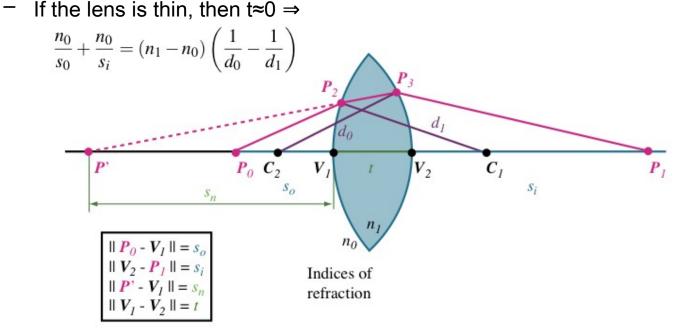
and if the image point is at  $\infty$ , l.e. if  $s_i=\infty$ , then:

- Object focal length:  $f = s_o = \frac{n_0}{d(n_1 n_0)}$ . Similarly, image focal length is obtained for  $s_0 = \infty$ .
  - Image focal length:  $f' = s_i = \frac{d n_1}{n_1 n_0}$

## Thin lenses

- However, lenses have a front and back surface
  - Spherical surfaces of radius d<sub>0</sub>
     and d<sub>1</sub>.
  - Analyzing front+back we have  $\frac{n_0}{s_0} + \frac{n_0}{s_i} = (n_1 n_0) \left( \frac{1}{d_0} \frac{1}{d_1} \right) + \frac{n_1 t}{(s_n t) s_n}$
- If lens is surrounded by air, then n<sub>0</sub>≈1 ⇒ *lens maker's formula*

$$\frac{1}{s_o} + \frac{1}{s_i} = (n_1 - 1) \left( \frac{1}{d_0} - \frac{1}{d_1} \right)$$



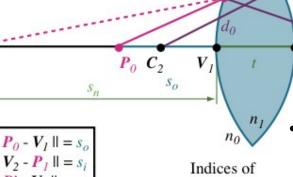
## Thin lenses

- If object distance s<sub>o</sub>=∞, the image distance becomes the image focal length s<sub>i</sub>=f<sub>i</sub>.
- Conversely, for points projected on an infinitely far away image plane, object focal length becomes s<sub>o</sub>=f<sub>o</sub>.

 But lens is thin, so we can set f<sub>i</sub>=f<sub>o</sub> and call it f

- So, we have  $\frac{1}{f} = (n_1 1) \left( \frac{1}{d_0} \frac{1}{d_1} \right)$
- Note: all rays at distance f in front of the lens, and passing through focal point, will be parallel after the lens (collimated light)
- Combining we find the *Gaussian lens* formula:

 $\frac{1}{f} = \frac{1}{s_o} + \frac{1}{s_i}$ Related to this: transverse (lateral) magnification:



refraction

which measures ratio of size of the image to the size of the object
Sign indicates whether it is upside down

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#### Thick lenses

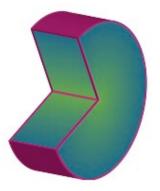
- Most lenses are not thin, they are thick
- We can think they are two spherical surfaces at a distance t
- Such lenses behave like the optical systems seen before, with 6 cardinal points
- If focal length is measured WRT principal planes, then

$$\frac{1}{f} = (n_1 - 1) \left( \frac{1}{d_0} - \frac{1}{d_1} + \frac{(n_1 - 1)t}{n_1 d_0 d_1} \right)$$

#### **Gradient lenses**

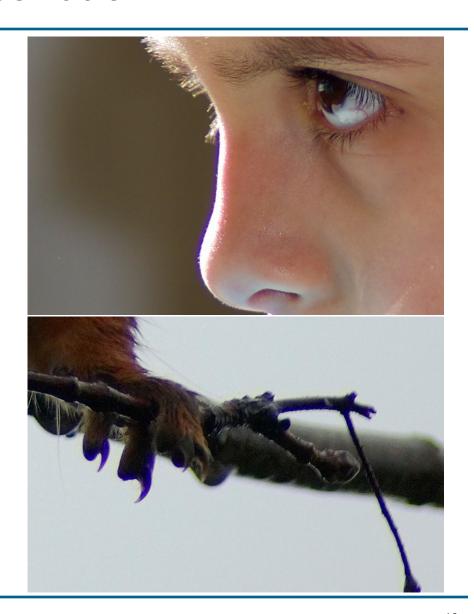
- Lens design characteristics considered up to now:
  - Dielectric material
  - Curvature of lens
- Placing more elements behind each other give additional flexibility
- However, one could build lenses having different refraction indexes at different places
- These are called gradient index lenses
- Obtained through immersing into salt solutions, which ionize and change refraction index

- In this case, one can use cylinders as lenses: these are called GRIN lenses
- Much more difficult to evaluate optically
- Raytracing may be used for this evaluation

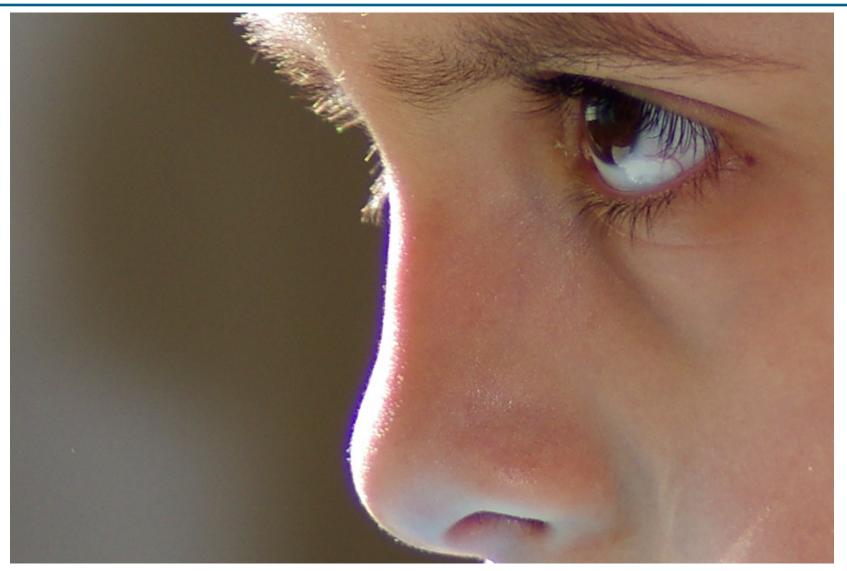


#### Lens aberration

- We did simple lens approximations, 1st order
- Lens computations must be higher order
- Often, ray tracing is used to evaluate lens design, which works well in theory
- However, deviations from ideal conditions occur: they are called aberrations
- These are of two types
  - Chromatic aberrations: the index of refraction is wavelength dependent: purple fringing
  - Rainbow-colour inaccuracies on edges



## Lens aberration



Curtesy Chem Kurmuk, pentaxforums.com

### Lens aberration

- Or one can have monochromatic aberrations
- Out of focus:
  - Spherical aberration
  - Astigmatism
  - Coma
- Warped image
  - Distortion
  - Petzval field distortion
- Reason for these aberrations:
  - We approximated sinus and cosinus linearly
    - sin=linear
    - cos=constant
  - Assumed paraxial rays

 This is a pretty rough approximation: we cut Taylor series to first term:

$$\sin(x) = \sum_{n=0}^{\infty} \frac{(-1)^n}{(2n+1)!} x^{2n+1}$$

$$\cos(x) = \sum_{n=0}^{\infty} \frac{(-1)^n}{(2n)!} x^{2n}.$$

- One could use higher order terms:
  - if one uses 2nd term, one obtains so-called *Third order theory*
  - Nicer, but higher complexity
- Aberrations here: due to Gaussian approximation

## **Spherical aberrations**

- For spherical lenses, we assumed that the ray has same length as path from object to image plane (on optical axis)
- If we keep term of 2nd degree, then equation

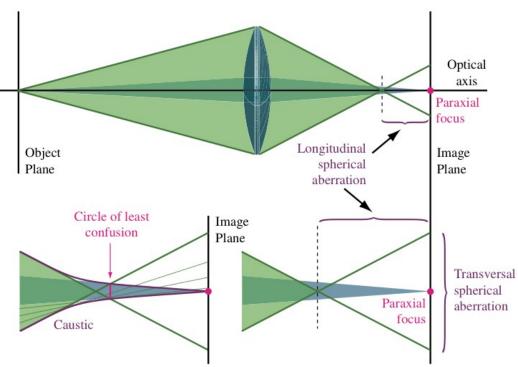
$$\frac{n_0}{s_o} + \frac{n_1}{s_i} = \frac{n_1 - n_0}{d}$$

becomes

$$\frac{n_0}{s_o} + \frac{n_1}{s_i} = \frac{n_1 - n_0}{d} + h^2 \left( \frac{n_0}{2s_o} \left( \frac{1}{s_0} + \frac{1}{d} \right)^2 + \frac{n_1}{2s_i} \left( \frac{1}{d} - \frac{1}{s_i} \right)^2 \right)$$

the extra term depends on h<sup>2</sup> with

 h=distance from point of lens where ray meets optical axis



- Light at lens border is focused nearer
- Defined as Spherical aberration

## Spherical aberrations

 Rays intersect optical axis over a length, not a point: longitudinal aberration.

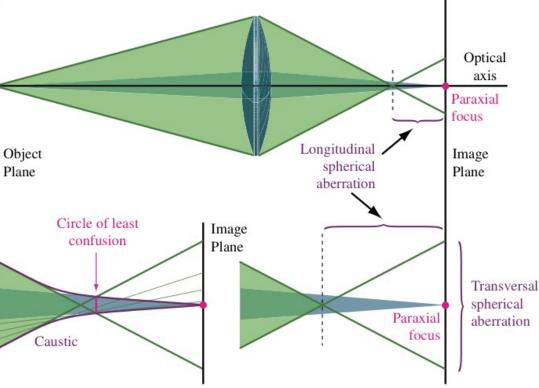
The rays will intersect image plane on a region: transversal aberration.

form curved convex hull: causticDiameter of projected spot

With spherical aberration, rays

smallest at circle of least confusion

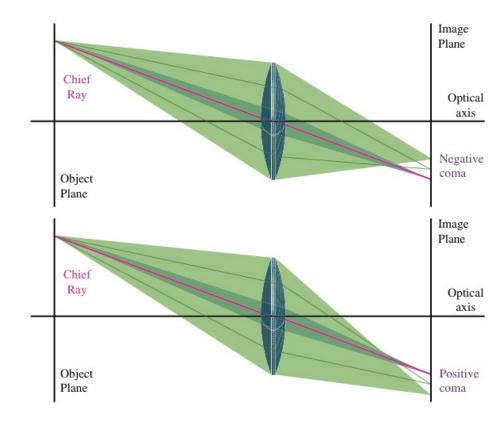




#### Coma

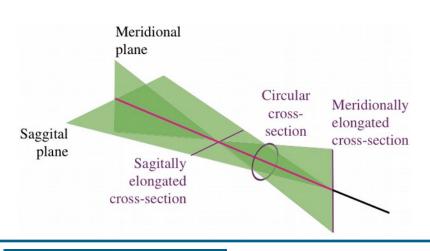
- Principal planes are well approximated only near the optical axis
- Further away, they are curved
- The effect is called coma:
  - Marginal ray focus farther than principal ray: positive coma
  - If closer, negative coma

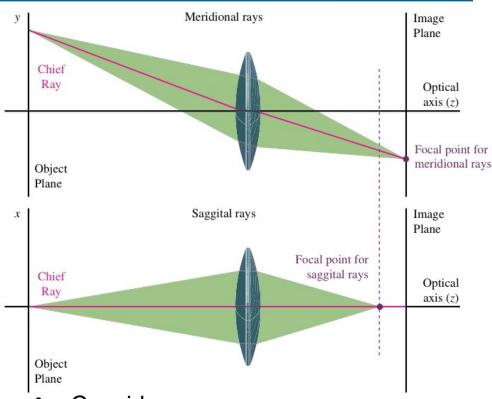




## **Astigmatic aberration**

- Occurs for off-axis object points
- Meridional plane was defined by object point and optical axis
- Chief ray lies in this plane but refracts at lens borders
- Sagittal plane:
  - − ⊥ meridional plane
  - Made by set of planar segments which intersect the chief ray

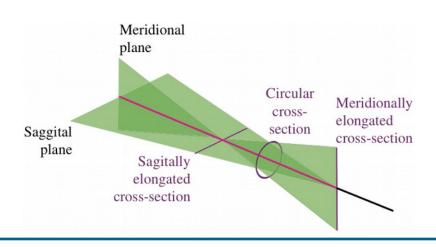


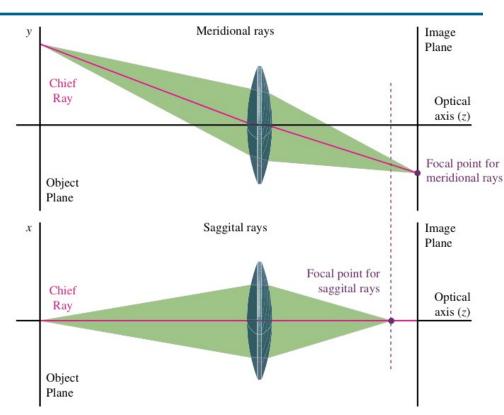


- Consider:
  - ray bundle in merid. plane
  - ray bundle in sagittal plane
  - Path length could be different

## **Astigmatic aberration**

- At sagittal focal point meridional rays will not have converged :
  - elongated focal point,
  - Elongated meridional focal point.
- For rays neither sagittal nor meridional, focal point will be in between the sagittal and meridional focal points.
- Somewhere between two focal points the cross-section of rays is circular





 When we have astigmatism, this circle is the place of sharpest focal point: circle of least confusion

#### Petzval field curvature

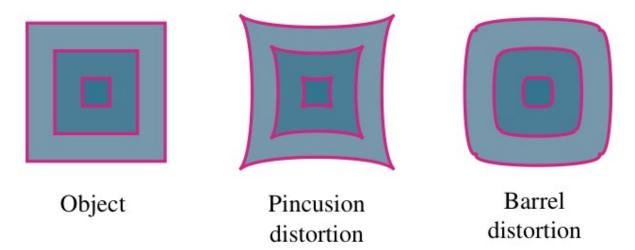
- For spherical lenses, object and image planes are not planar, but:
  - Positive lens: curve inwards
  - Negative lens: curve outwards
  - Petzval field curvature
- If a flat image plane is used,
   I.e. on a sensor, the image will only be sharp on optical axis
- One can correct this by combining positive and negative lenses

 Example: correct inward curvature of positive lens with negative length near focal point of positive lens: field flattener.

#### **Distortion**

- Distortion is due to lateral magnification of lens:
  - Lateral magnification is not constant as assumed before
- Pincushion distortion:
   lateral magnification increases with distance to optical axis
  - Usually positive lenses generate it
- Barrel distortion: lat. mag. decreases with distance to optical axis
  - Usually negative lenses generate it

5/21/19



#### **Cromatic aberrations**

- Materials have wavelength-dependent refraction index, which influences focal length
  - Thus focal length a white light beam lies closer for blue rays than to red light
  - Distance between these two points on optical axis is the axial chromatic aberration
  - These rays will hit image at different position (*lateral chromatic aberration*)
  - Can be corrected by using thin lenses with different refractive indexes (thin achromatic doublets):
    - If d is the distance between the lenses, wavelength dep. focal length is  $f_1(\lambda)$  and  $f_2(\lambda)$ , refraction indexes  $n_1(\lambda)$  and  $n_2(\lambda)$  wavelength dependent focal length  $f(\lambda)$  is then given by

$$\frac{1}{f(\lambda)} = \frac{1}{f_1(\lambda)} + \frac{1}{f_2(\lambda)} - \frac{d}{f_1(\lambda) f_2(\lambda)}$$

## **Cromatic aberrations**

 If the index of refraction of surrounding medium is 1, then

$$\frac{1}{f_1(\lambda)} = k_1 (n_1(\lambda) - 1)$$
$$\frac{1}{f_2(\lambda)} = k_2 (n_2(\lambda) - 1)$$

is wavelength dependent focal length. Here, we replaced factor depending on front and back radius with constants k<sub>1</sub>,k<sub>2</sub>.

Substituting,

$$\frac{1}{f(\lambda)} = k_1 (n_1(\lambda) - 1) + k_2 (n_2(\lambda) - 1) - \frac{d}{\frac{1}{k_1 (n_1(\lambda) - 1)}} \frac{1}{k_2 (n_2(\lambda) - 1)}$$

• For the two focal lengths  $f(\lambda_R)$  and  $f(\lambda_B)$  to be equal, one must place lenses at a distance given by solving for d:

$$d = \frac{1}{k_1 k_2} \frac{k_1 (n_1(\lambda_B) - n_1(\lambda_R)) + k_2 (n_2(\lambda_B) - n_2(\lambda_R))}{(n_1(\lambda_B) - 1) (n_2(\lambda_B) - 1) - (n_1(\lambda_R) - 1) (n_2(\lambda_R) - 1)}$$

If lenses touch, d=0:

$$\frac{k_1}{k_2} = -\frac{n_2(\lambda_B) - n_2(\lambda_R)}{n_1(\lambda_B) - n_1(\lambda_R)}$$

 Now we can have focal length of yellow light (λ<sub>Y</sub>≈λ<sub>R</sub>+λ<sub>B</sub>)/2)

$$\frac{1}{f_1(\lambda_Y)} = k_1 (n_1(\lambda_Y) - 1)$$
$$\frac{1}{f_2(\lambda_Y)} = k_2 (n_2(\lambda_Y) - 1)$$

which is

$$\frac{k_1}{k_2} = \frac{n_2(\lambda_Y) - 1}{n_1(\lambda_Y) - 1} \frac{f_2(\lambda_Y)}{f_1(\lambda_Y)}$$

## **Chromatic aberrations**

So, we have:

$$\frac{f_2(\lambda_Y)}{f_1(\lambda_Y)} = \frac{(n_2(\lambda_B) - n_2(\lambda_R)) / (n_2(\lambda_Y) - 1)}{(n_1(\lambda_B) - n_1(\lambda_R)) / (n_1(\lambda_Y) - 1)} = \frac{w_2}{w_1}$$
 where  $w_1, w_2$  are the dispersive powers associated with the refraction indexes  $n_1, n_2$ .

 Take the standardized
 Fraunhofer spectral lines F,D,C
 and the wavelengths

$$\lambda_F = 486.1 \text{ nm}$$
 $\lambda_D = 589.2 \text{ nm}$ 
 $\lambda_C = 656.3 \text{ nm}$ 

We can now define dispersive power of an optical material w:

$$V = \frac{n_D - 1}{n_F - n_C}$$

where V=1/w and is called *Abbe number*, or *dispersive index*.

- Here,  $n_D = n(ID)$ , .
- For lenses, it is desirable to have materials with low dispersion, or high Abbe numbers.

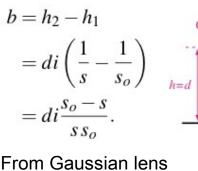
## Blur circle

- An ideal optical system would image a point source onto a single point on the image plane.
- Due to aberrations: blurred shape on the image plane
  - Can be approximated as a circle: its radius can be approximated as follows:
  - Place object at same height h of lens aperture height
  - Radius b of blur circle is:

- Rewriting:  $i = \frac{sf}{s-f}$
- Thus:

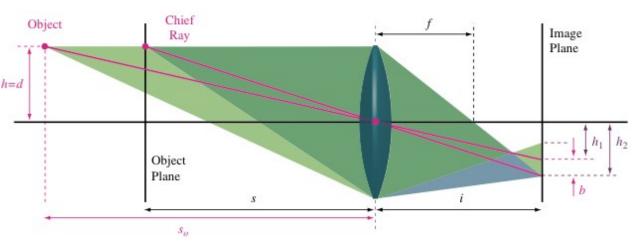
$$b = \frac{sfd}{s - f} \frac{s_o - s}{s s_o} = \frac{pfd}{s - f}$$

where  $p(s_0-s)/ss_0$  can be seen as percentage focus error.



 From Gaussian lens formula

$$\frac{1}{f} = \frac{1}{s} + \frac{1}{i}$$



## Depth of field



- Points on image plane have maximum sharpness.
- Objects are not all on image plane: some before, some after.
- There is a region in which points are focused reasonably sharp: depth of field.
- Sharpness depends on:
  - Size of image plane
  - Sensor resolution
  - Image reproduction size
  - Angular resolution of human visual system

## Depth of field

- If we can make sure that a circle of radius b leads to an image that in the reproduction appears as a single point, then, assuming
  - Camera focused at distance  $s_o$
  - − Blur circle smaller aperture ( $b \ll d$ )
- Then distance of lens to nearest point  $s_{near}$  of acceptable focus is

$$s_{\text{near}} = \frac{s_o f}{f + \frac{b}{d} (s_0 - f)}$$

and for farthest point it is

$$s_{\text{far}} = \frac{s_o f}{f - \frac{b}{d} (s_o - f)}$$

• The global depth of field is then  $s_{far}$ - $s_{near}$ 

 Far plane becomes infinite when denominator goes to 0, i.e. when

$$f = \frac{b}{d} (s_o - f)$$

solve for s<sub>o</sub>: hyperfocal distance

$$s_o = f\left(\frac{d}{b} + 1\right)$$

which corresponds to the near plane:

$$s_{\text{near}} = \frac{f}{2} \left( \frac{d}{b} + 1 \right)$$

so, if the camera fucused on the hyperfocal plane, all objects between the near plane and infinity will be in focus

• Notice that the aperture *d* affects the depth of field!

## **Depth of field**





f/3.2 f/16